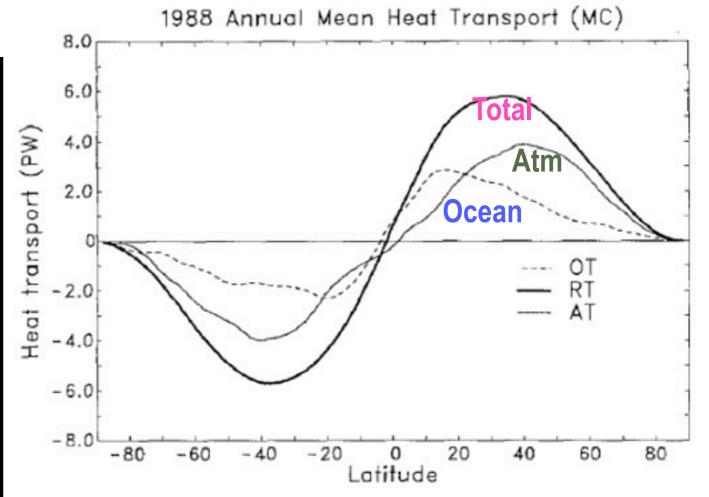


Solar radiation Incoming/Outgoing

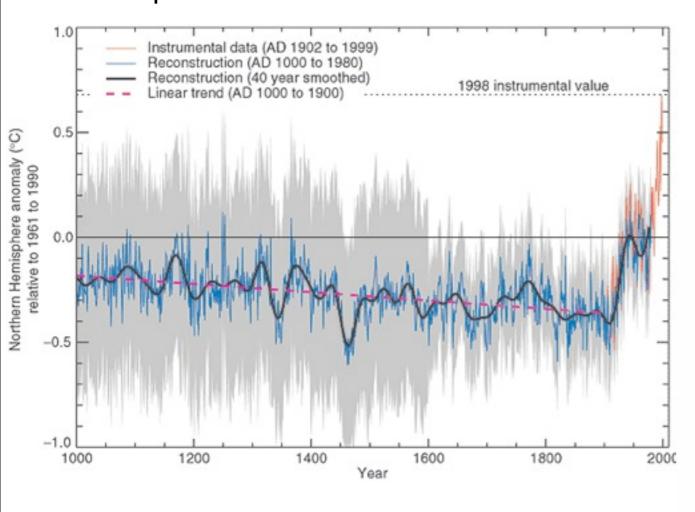
measured at top of atmosphere

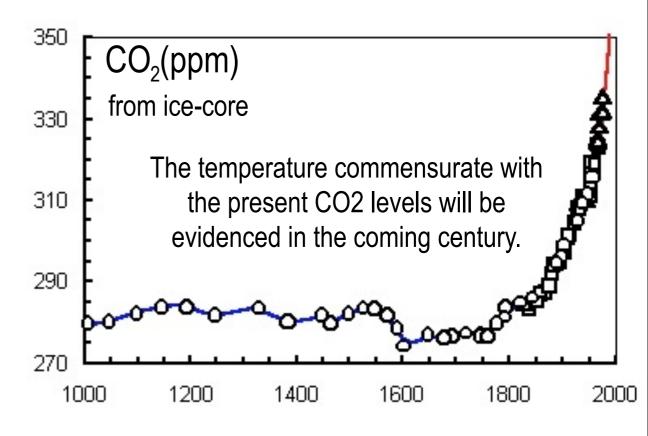


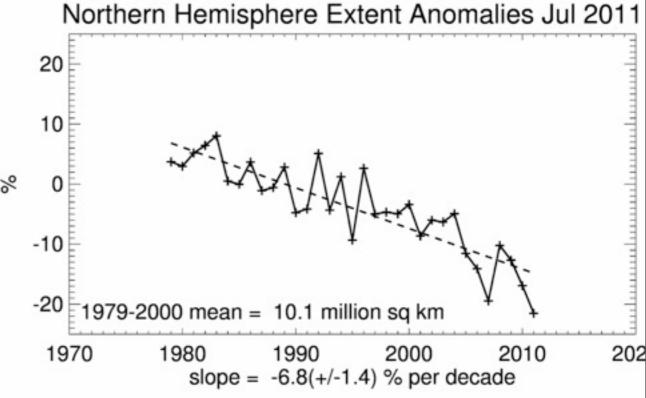
Mean Meridional Heat Transport Trenberth and Solomon (1994)

We are presently experiencing in one of the largest and most abrupt perturbations to climate and ecology

Temperature







Millennial Northern Hemisphere (NH) temperature reconand instrumental data (red) from AD 1000 to 1999, adapted from Mann et al. (1999).

IPCC report, 2001

Global Temperature (Hansen et al., 2006)

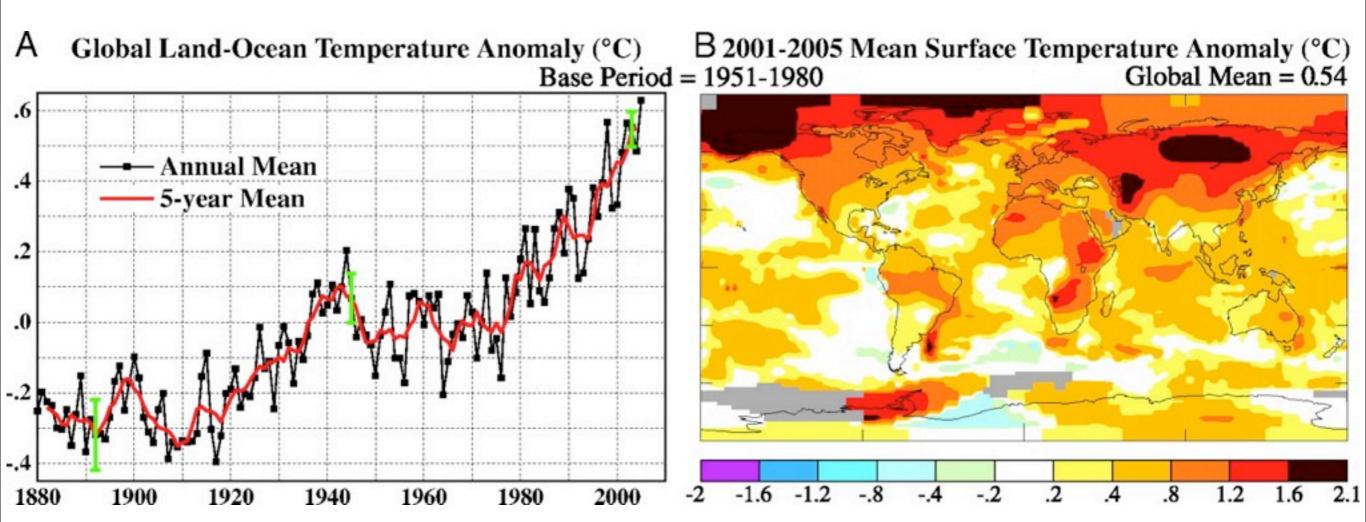


Fig. 1. Surface temperature anomalies relative to 1951-1980 from surface air measurements at meteorological stations and ship and satellite SST measurements. (A) Global annual mean anomalies. (B) Temperature anomaly for the first half decade of the 21st century.

What generates the complex pattern of flow?

Density variations

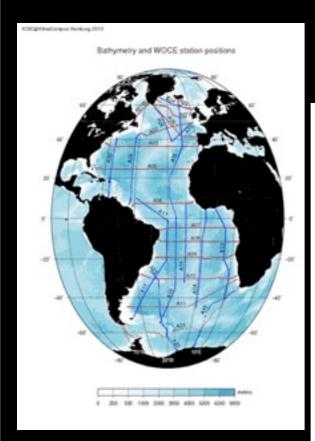
Effects of rotation large scale motion- geostrophy

Mechanical forcing by wind

Fronts Lateral density variation

Baroclinic instability - eddy formation, mixing

Vertical motion

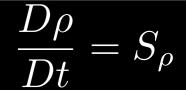


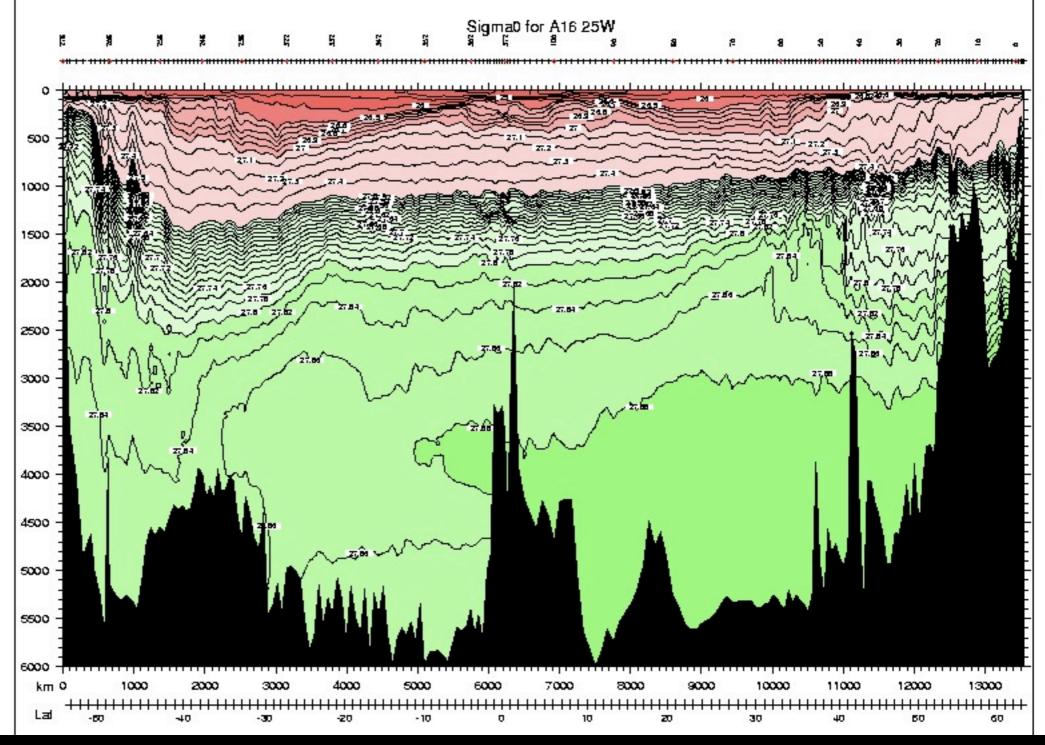
$$N^2 = -\frac{g}{\rho} \frac{\partial \rho}{\partial z}$$

Buoyancy frequency

1024-1029 kg/m3

WOCE - Atlantic section potential density





Internal gravity waves

Breaking leads to diapycnal mixing

Dense fluid collapse

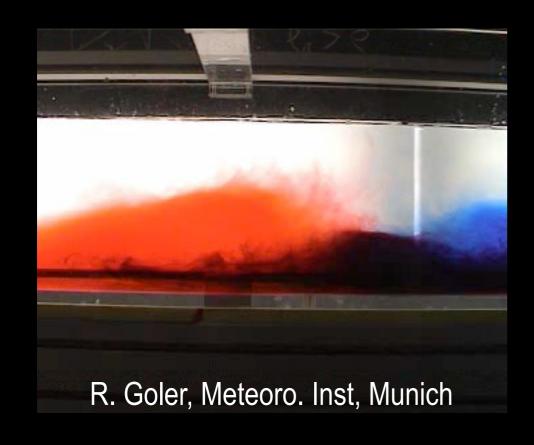
with rotation

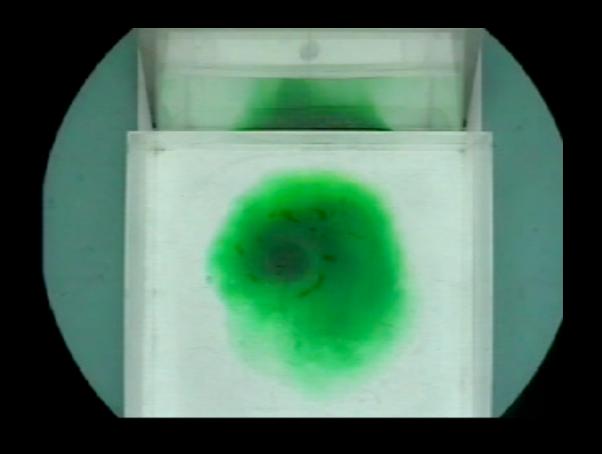


MIT Synoptic Meteero. Lab

Dense fluid collapse

with rotation





MIT Synoptic Meteero. Lab

Equations of motion

in a rotating frame
$$\left(\frac{\mathrm{d}r}{\mathrm{d}t} \right)_{I} = \left(\frac{\mathrm{d}r}{\mathrm{d}t} \right)_{R} + \Omega \times r$$

$$v_I = v_R + \Omega \times r$$
.

$$\left(\frac{\mathrm{d}\boldsymbol{v}_R}{\mathrm{d}t}\right)_{\!\!R} = \left(\frac{\mathrm{d}\boldsymbol{v}_I}{\mathrm{d}t}\right)_{\!\!I} - 2\boldsymbol{\varOmega} \times \boldsymbol{v}_R - \boldsymbol{\varOmega} \times (\boldsymbol{\varOmega} \times \boldsymbol{r})$$
Coriolis accn

$$\frac{D\mathbf{u}}{Dt} + \rho^{-1}\nabla p + 2\Omega \times \mathbf{u} + \nabla \phi = \nabla(\nu \nabla u)$$

$$\frac{U^2}{L} \qquad \frac{P}{\rho L} \qquad \Omega U \qquad g \qquad \frac{\nu U}{L^2}$$

Effects of rotation We view the earth in the rotating frame

Effects of rotation We view the earth in the rotating frame



Horizontal Large Scale Dynamics

$$Du/Dt + \left(\frac{P}{\rho U^2}\right)p_x - \left(\frac{\Omega L}{U}\right)fv = \left(\frac{\nu}{UL}\right)\nabla^2 u$$

For
$$U = 0.1 \ m/s$$
, $L = 10^5 \ m$, $\Omega = 10^{-4}/s$

Rossby number
$$R_o = \frac{U}{\Omega L} << 1$$
 $Re = UL/\nu \sim 10^{10}$

$$Re = UL/\nu \sim 10^{10}$$

$$b \equiv 2\Omega \cos \phi$$

$$2\Omega \qquad f \equiv 2\Omega \sin \phi$$

 $2\Omega \times \mathbf{u} = (-fv + bw, fu, bu)$

$$Du/Dt + Ro^{-1}(p_x - fv + Ro\delta bw) = F^x$$
$$Dv/Dt + Ro^{-1}(p_y + fu) = F^y$$

Geostrophy

$$fv = p_x, \quad fu = -p_y$$

$$u_x + v_y + w_z = 0$$

In the Vertical

$$\left(\frac{U}{L}\right)(u_x + v_y) + \left(\frac{W}{D}\right)w_z = 0$$
$$\delta = D/L << 1$$

$$Ro = \frac{U}{\Omega L}$$

$$w_z = -Ro^{-1}(u_x + v_y)$$

$$W \sim Ro \delta U$$

$$\frac{Dw}{Dt} + \frac{1}{Ro^2\delta} \left(\frac{1}{\rho} p_z + g - \delta bu \right) = 0$$

Hydrostatic balance

$$\frac{\partial p}{\partial z} = -\rho g$$

At the sea surface

$$v_g = \frac{1}{f}gh_x, \quad u_g = -\frac{1}{f}gh_y$$

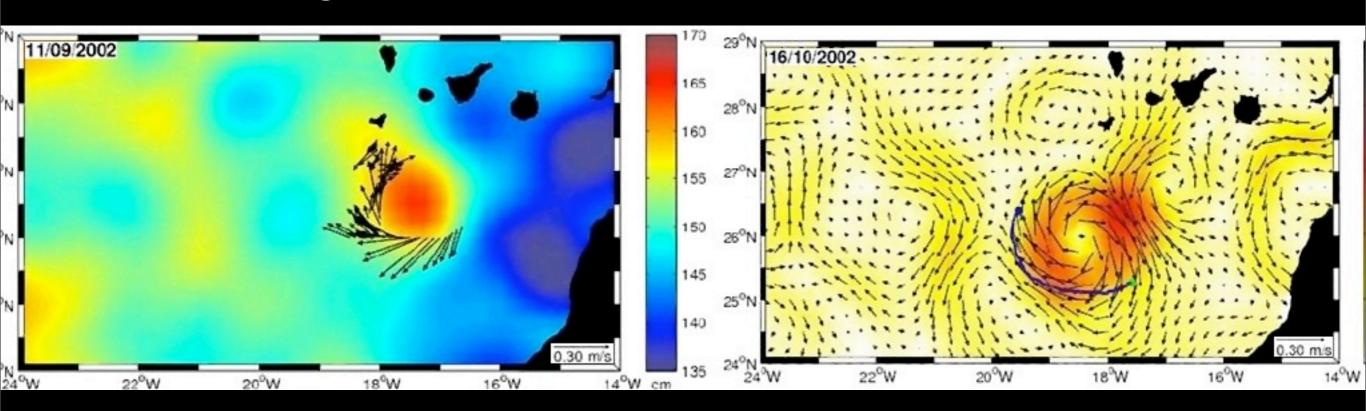
Hydrostatic Pressure gradient

$$p_x = gh_x + r_x$$

$$r_x = \frac{\partial}{\partial x} \int_z^h \rho dz$$

Sea surface height

Geostrophic surface velocity



In the Boundary Layer

Ekman transport

Incompressibility

$$2\Omega \times \mathbf{u} = \frac{\partial}{\partial z} \left(K_v \frac{\partial u}{\partial z} \right) \qquad K_v \frac{\partial u}{\partial z} = \tau / \rho$$

$$\int U \qquad \frac{K_v U}{H^2}$$

w e from Williams

$$E_k = \frac{K_v}{fH_E^2} = O(1)$$
 $H_E = (f/K_v)$

$$M_E = \int_{\mathrm{Ek}} u_E \, \mathrm{d}z$$
 $M_E = -\frac{1}{f} \mathbf{k} \times \widetilde{\boldsymbol{ au}}_t$ $\nabla \cdot M_E = w_E = \nabla \times \boldsymbol{ au}$ AND $\nabla \cdot M_E = w_E = \nabla \cdot \mathbf{x}$ $\nabla \cdot$

from L.D. Stott

The wind-driven circulation

Surface boundary condition

$$\nu \frac{\partial \mathbf{u}}{\partial z} = \tau / \rho$$

Stommel model

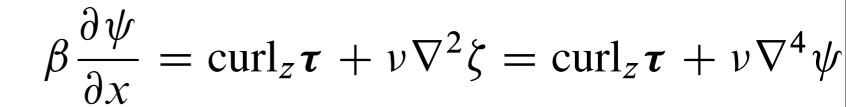
$$2\Omega \times \mathbf{u} = \frac{1}{\rho} \frac{\partial \tau}{\partial z}$$

Take the curl and integrate over depth of BL

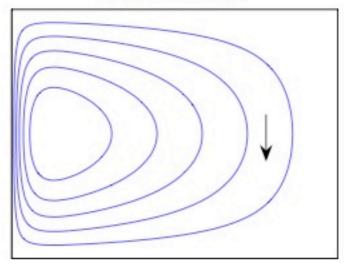
$$\int f \nabla_z \cdot \boldsymbol{u} \, dz + \frac{\partial f}{\partial v} \int v \, dz = \operatorname{curl}_z (\tilde{\boldsymbol{\tau}}_t - \tilde{\boldsymbol{\tau}}_b)$$

Sverdrup balance: Vertically integrated

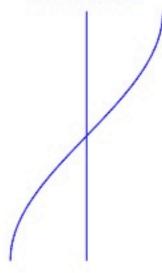
$$\beta \overline{v} \approx \operatorname{curl}_z \widetilde{\boldsymbol{\tau}}$$

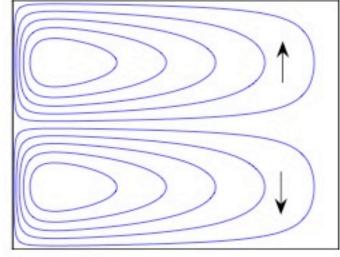


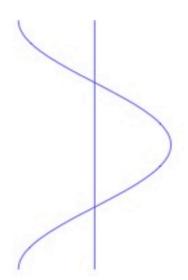
Streamfunction



Wind Stress

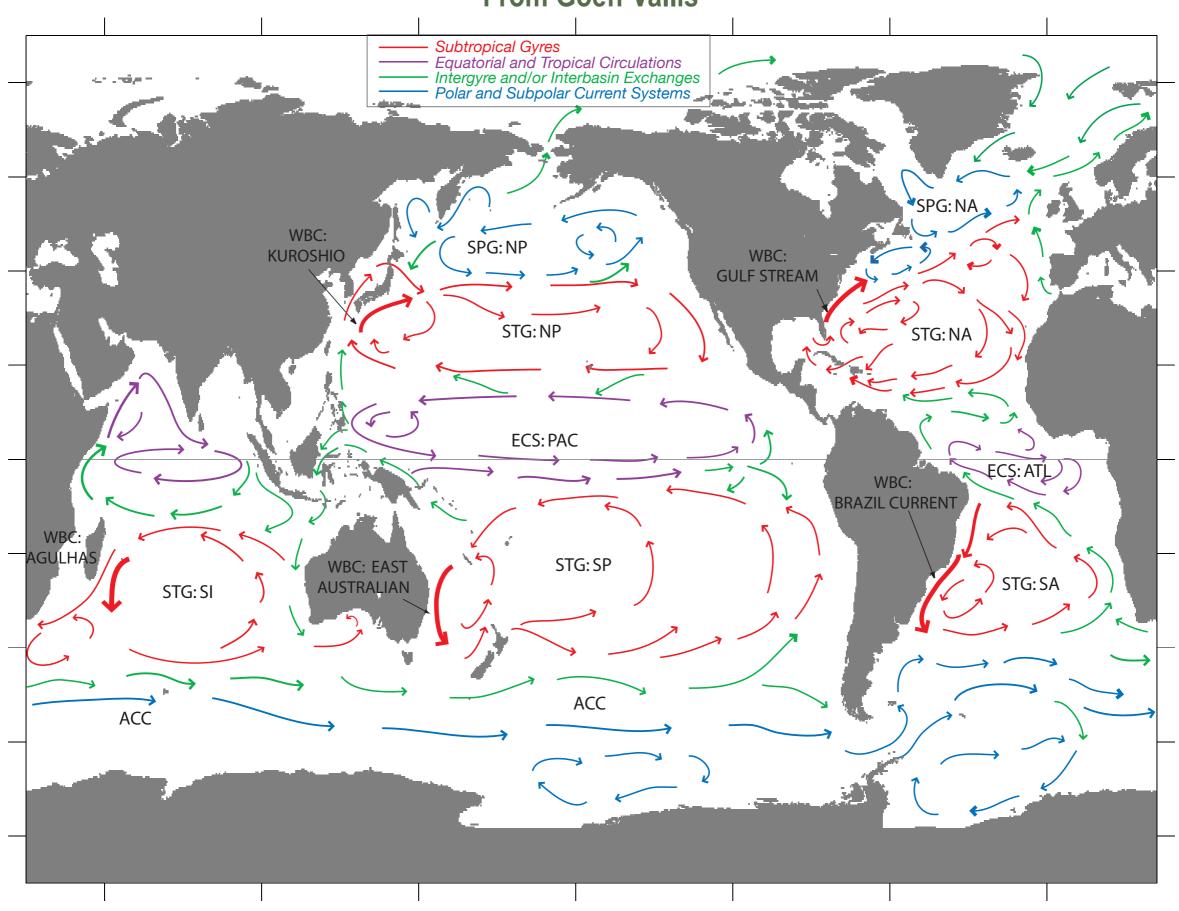






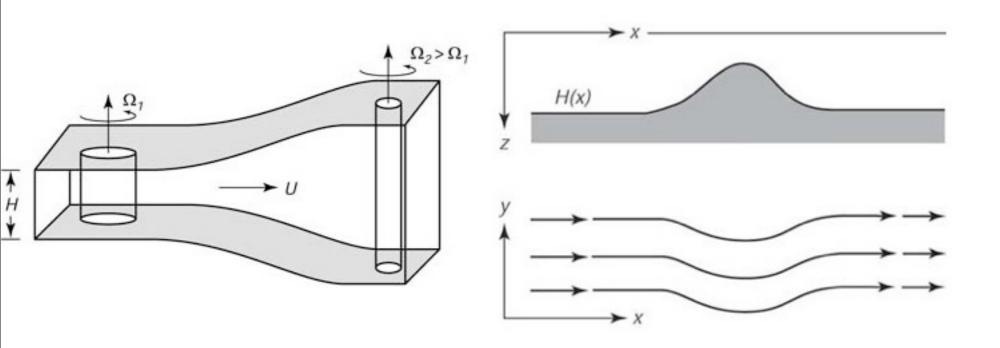
The wind-driven gyres

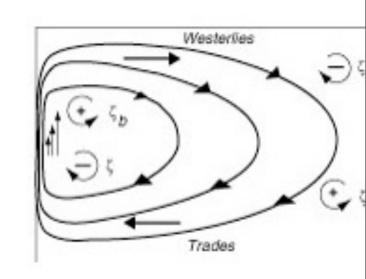
From Goeff Vallis



Conservation of Potential Vorticity

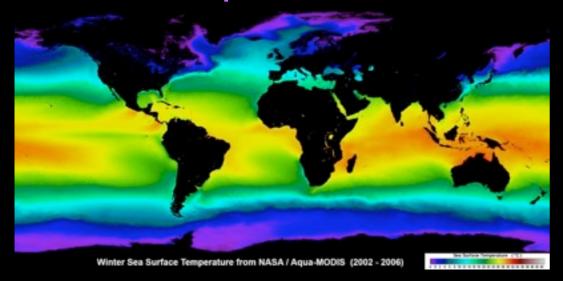
$$\frac{d}{dt}\left(\frac{f+\zeta}{h}\right) = 0, \qquad \zeta = v_x - u_y$$

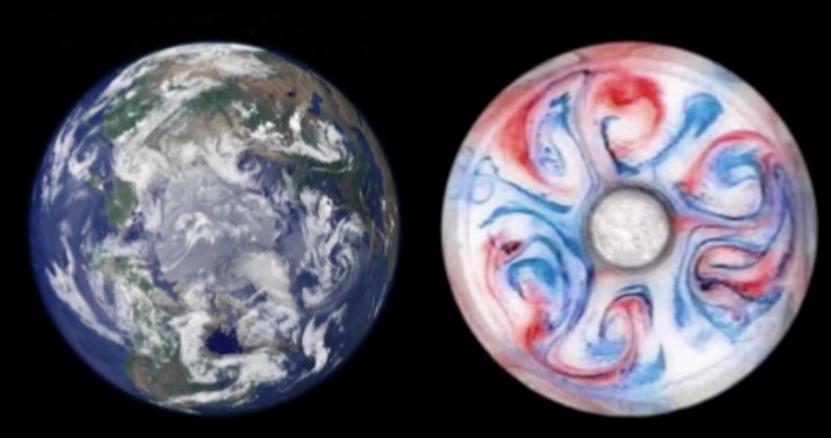


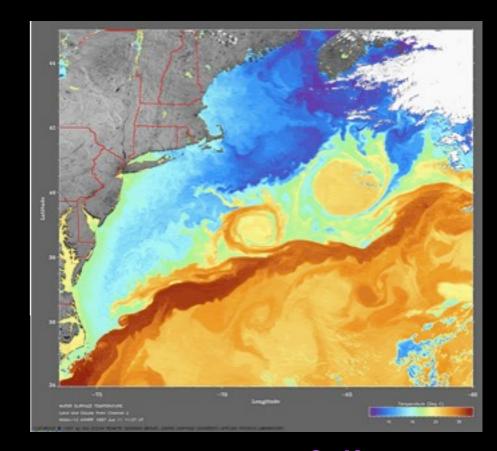


Lateral variations in density

Sea Surface Temperature from Satallite

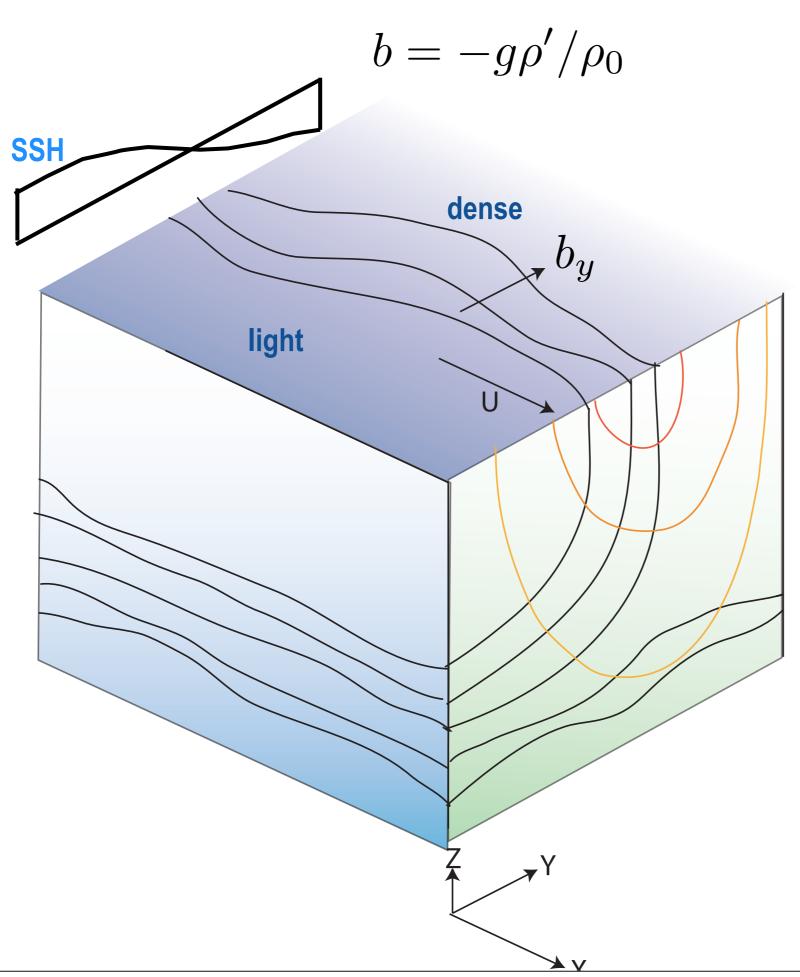






Gulf stream

from Marshall and Plumb 2006



Hydrostatic balance

$$p_z = -\rho g$$

Geostrophic balance

$$fu = -p_y$$

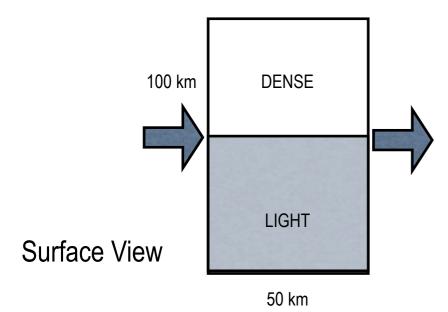
Thermal wind balance

$$fu_z = g\rho_y$$

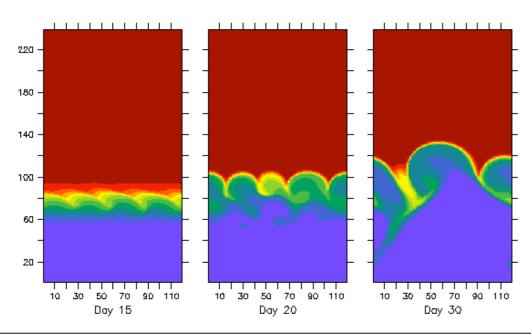
Numerical Experiments

Periodic Channel

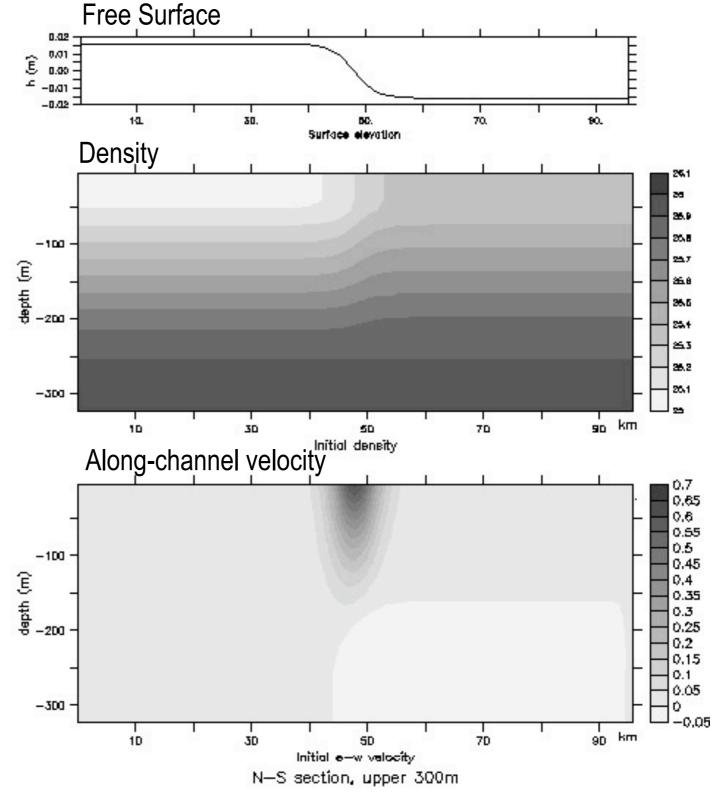
Initial Conditions



Varying resolution 1000-250m Wind / no wind Hydrostatic/Nonhydrostatic







Instability of a Front

Dispersion/ cross-front exchange

Cold -blue Warm - red

Vorticity / f

200 km

z (m)

y (km)

SECTIONAL VIEW

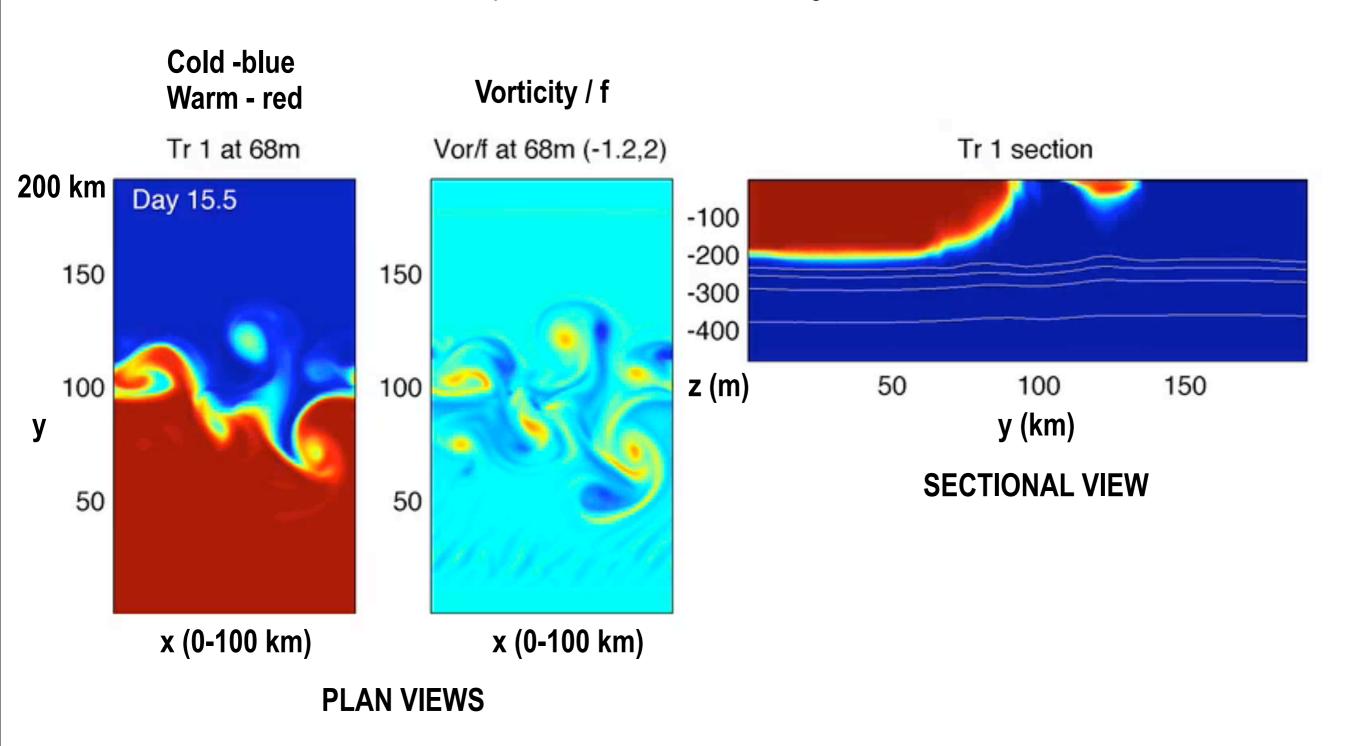
x (0-100 km)

x (0-100 km)

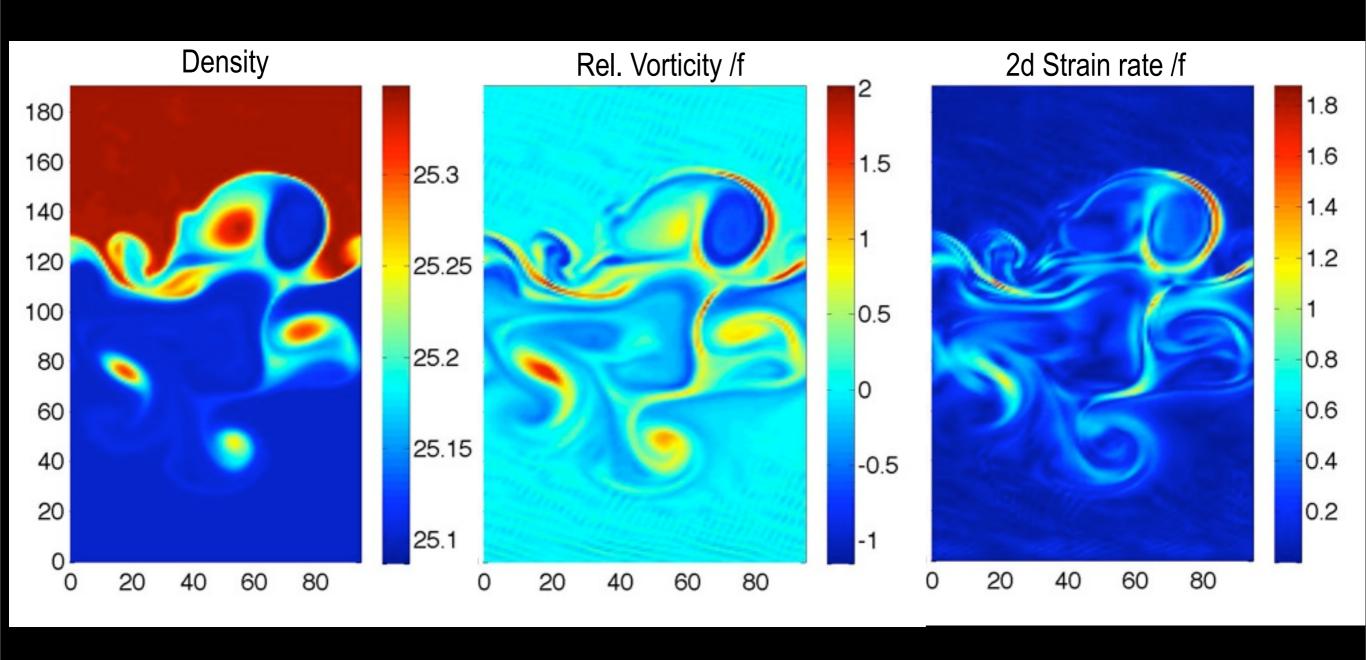
PLAN VIEWS

Instability of a Front

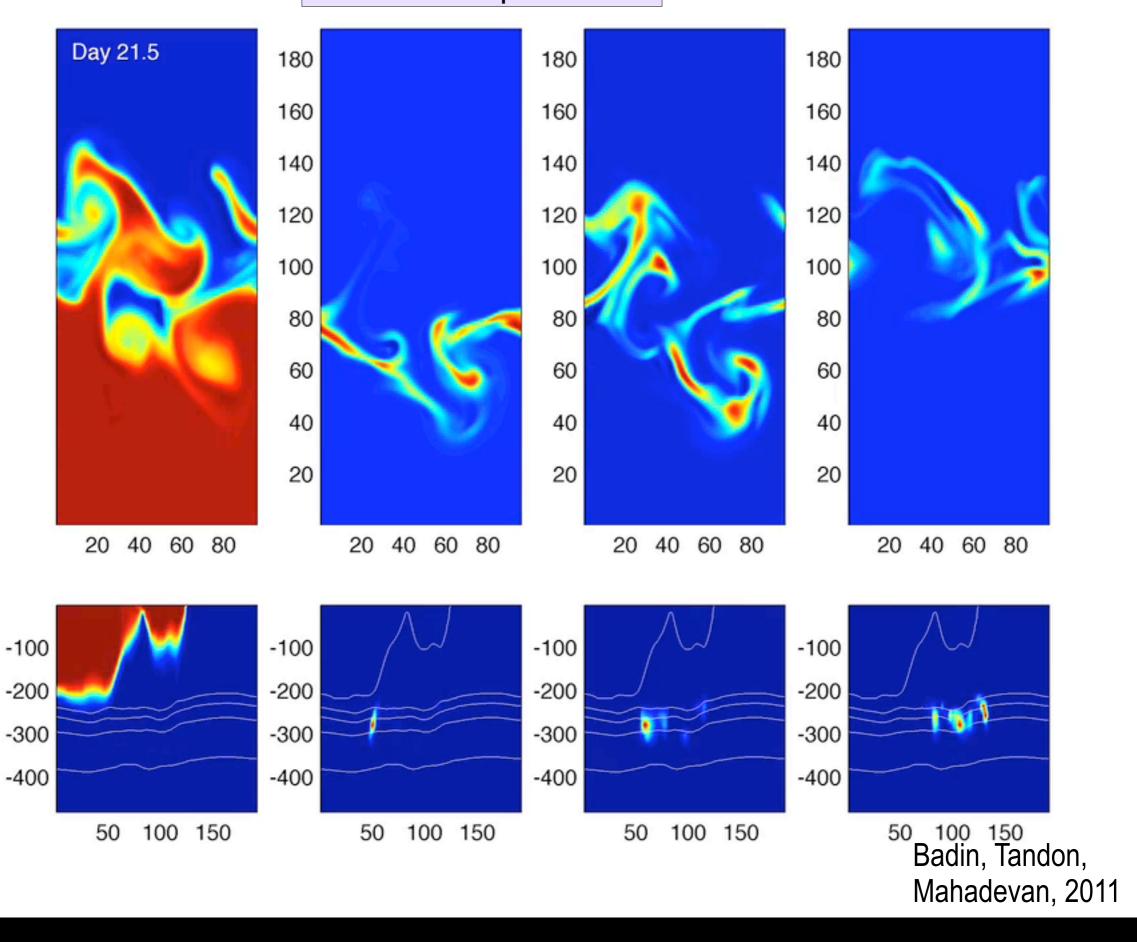
Dispersion/ cross-front exchange



Front - forms eddies and filaments



Tracer dispersion



Tracer variance

First Moment

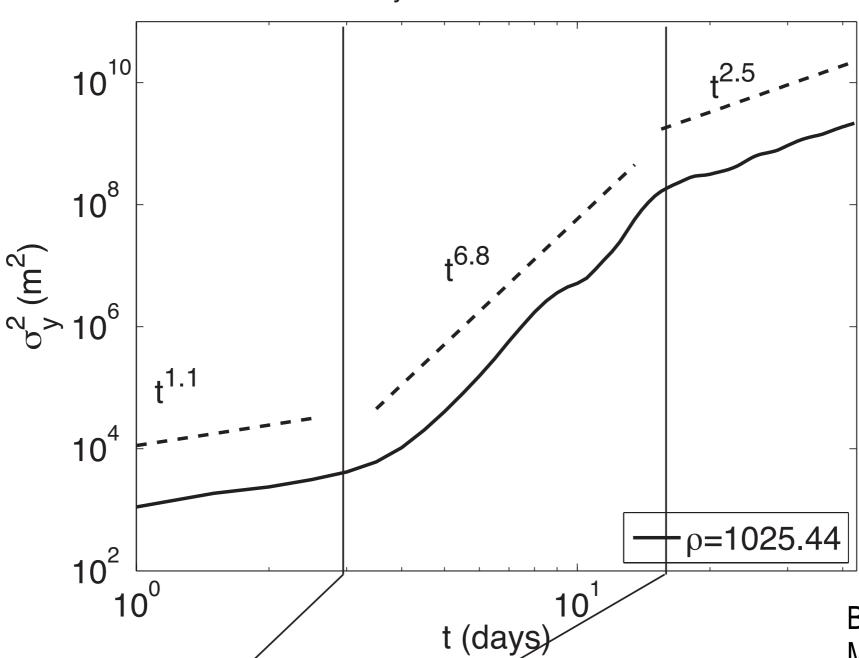
 $\mathbf{d}_{\mathbf{y}} = \frac{\langle \mathbf{y} \mathbf{C} \rangle}{\langle \mathbf{C} \rangle}$

Position of center of mass

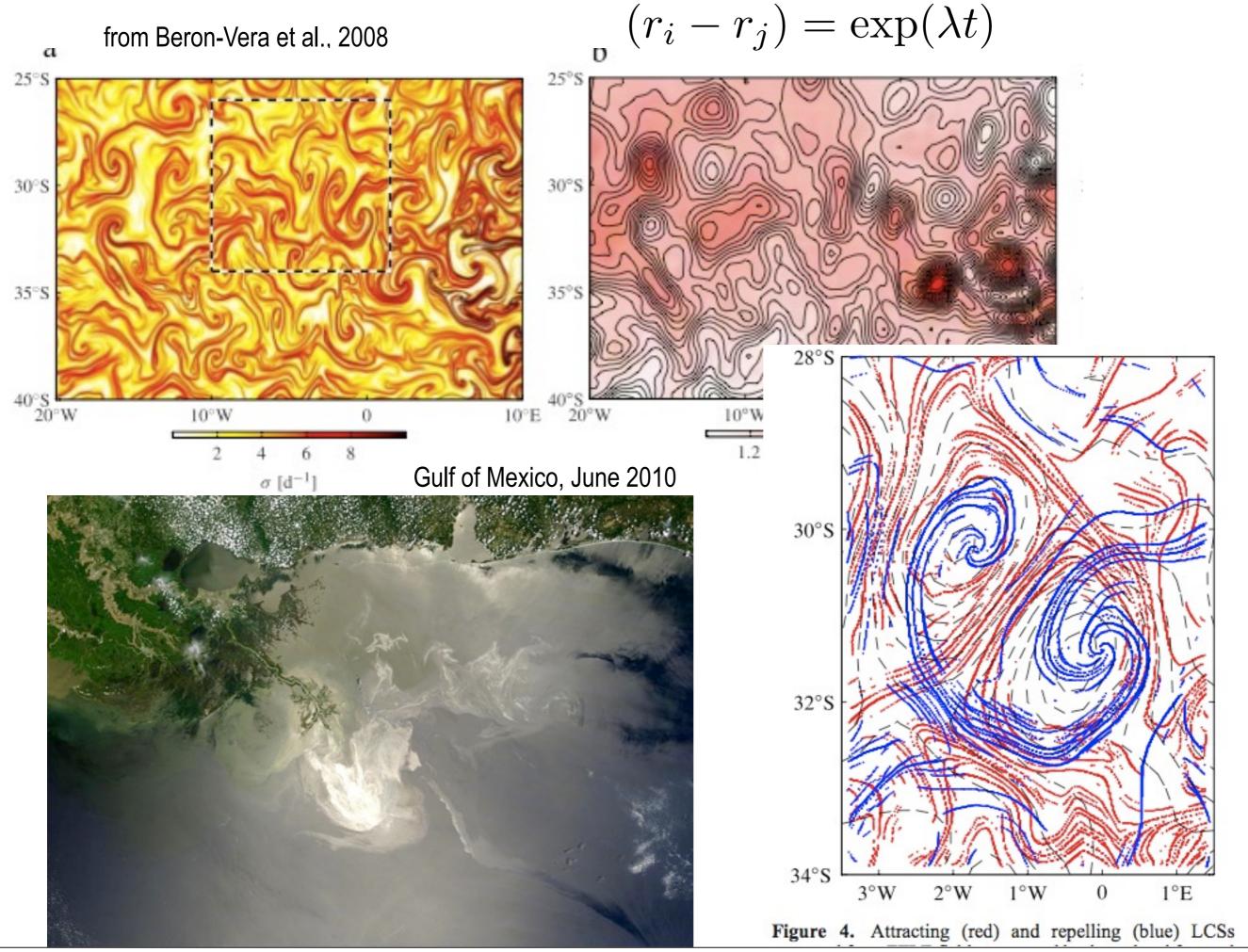
Dispersion ~Variance

 $\sigma_{y}^{2} = \frac{\langle y^{2}C \rangle}{\langle C \rangle} - d_{y}^{2}$





Badin, Tandon, Mahadevan, 2011



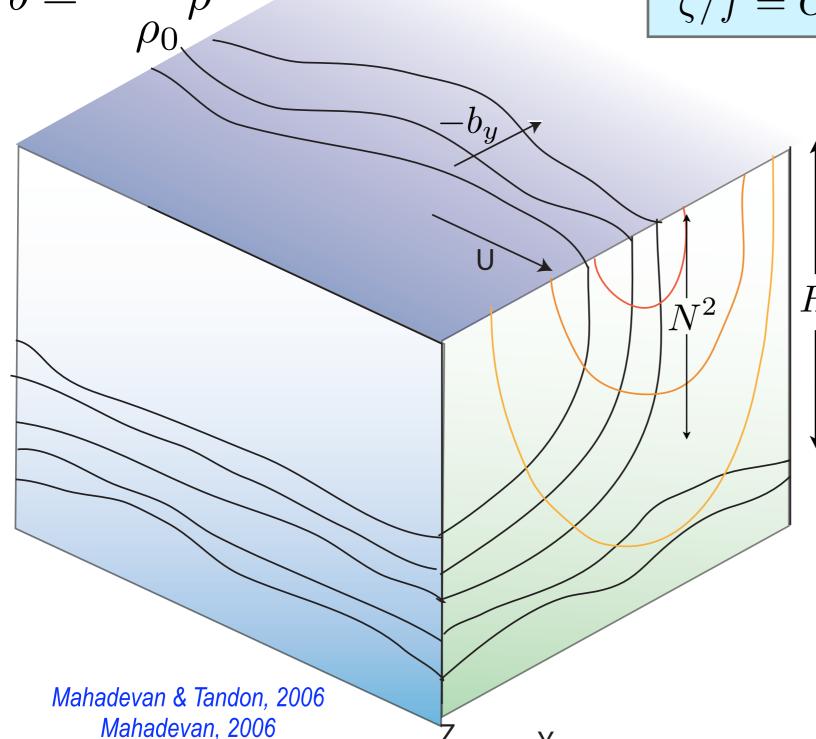
Fronts - Lateral gradients in density

Vertical transport

$$b = -\frac{g}{2\sigma}\rho'$$

$$\zeta/f = O(1)$$

$$\zeta/f = O(1), \quad Ro = U/fL = O(1)$$



Vertical velocity

$$W \sim Ro \delta U = \delta U$$

where $\delta = H/L = f/N$

$$\frac{\partial}{\partial y} \quad \frac{Db}{Dt} = 0$$

Frontogenesis

$$\frac{Db_y}{Dt} = -u_y b_y - v_y b_y$$

Review on submesoscale processes:

Thomas, Tandon, Mahadevan, 2008

Modeling

p = Hydrostatic pressure

q = Nonhydrostatic pressure

 $P = p + \delta q$

$$D_t u + Ro^{-1}(p_x + \delta q_x - fv + Ro \delta bw) = F^x$$

$$D_t v + Ro^{-1}(p_y + \delta q_y + fu) = F^y$$

$$D_t w + Ro^{-2} \delta^{-2} (\rho^{-1} p_z + g + \delta q_z - \delta bu) = F^{\frac{1}{2}}$$

$$u_x + v_y + Ro \ w_z = 0$$

Hydrostatic $\delta \rightarrow 0$

$$p_z + \rho g = 0$$

$$0 \qquad w_z = -Ro^{-1}(u_x + v_y)$$

 $W \sim Ro \delta U$

 $b \equiv 2\Omega \cos \phi$

 $\equiv 2\Omega \sin \phi$

Free-surface height ...and density => p

$$h_t + \partial_x \int_{z_b}^h u dz + \partial_y \int_{z_b}^h v dz = 0$$

Nonhydrostatic (δ does not \rightarrow 0)

$$D_t w + Ro^{-2} \delta^{-1} (q_z - bu) = F^z$$

Well-posed with open boundaries

Mahadevan et al., 1996a,b, Mahadevan & Archer 1998

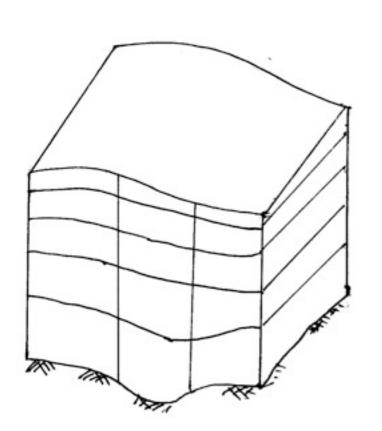
Nonhydrostatic Model

3-D pressure field to be determined

Using incompressibility

$$q_{xx} + q_{yy} + \delta^{-2}q_{zz} = F$$

Discretized ...
$$(q_{i+1} - 2q_i + q_{i-1}) + (q_{j+1} - 2q_j + q_{j-1}) + \delta^{-2}(q_{k+1} - 2q_k + q_{k-1}) = F$$



$$\begin{bmatrix} 2/\delta^2 & \delta^{-2} & \dots & 1 & 1 & \dots \\ 1/\delta^2 & 2/\delta^2 & 1/\delta^2 & \dots & 1 & 1 \\ \dots & 1/\delta^2 & 2/\delta^2 & 1/\delta^2 & \dots & 1 \\ \dots & \dots & \dots & \dots & \dots \\ 1 & 1 & \dots & 1/\delta^2 & 2/\delta^2 & 1/\delta^2 \\ \dots & 1 & 1 & \dots & 1/\delta^2 & 2/\delta^2 \end{bmatrix} \begin{bmatrix} \dots \\ \dots \\ q_{ijk} \\ \dots \\ \dots \\ \dots \end{bmatrix} = \begin{bmatrix} \dots \\ \dots \\ F_{ijk} \\ \dots \\ \dots \\ \dots \\ \dots \end{bmatrix}$$

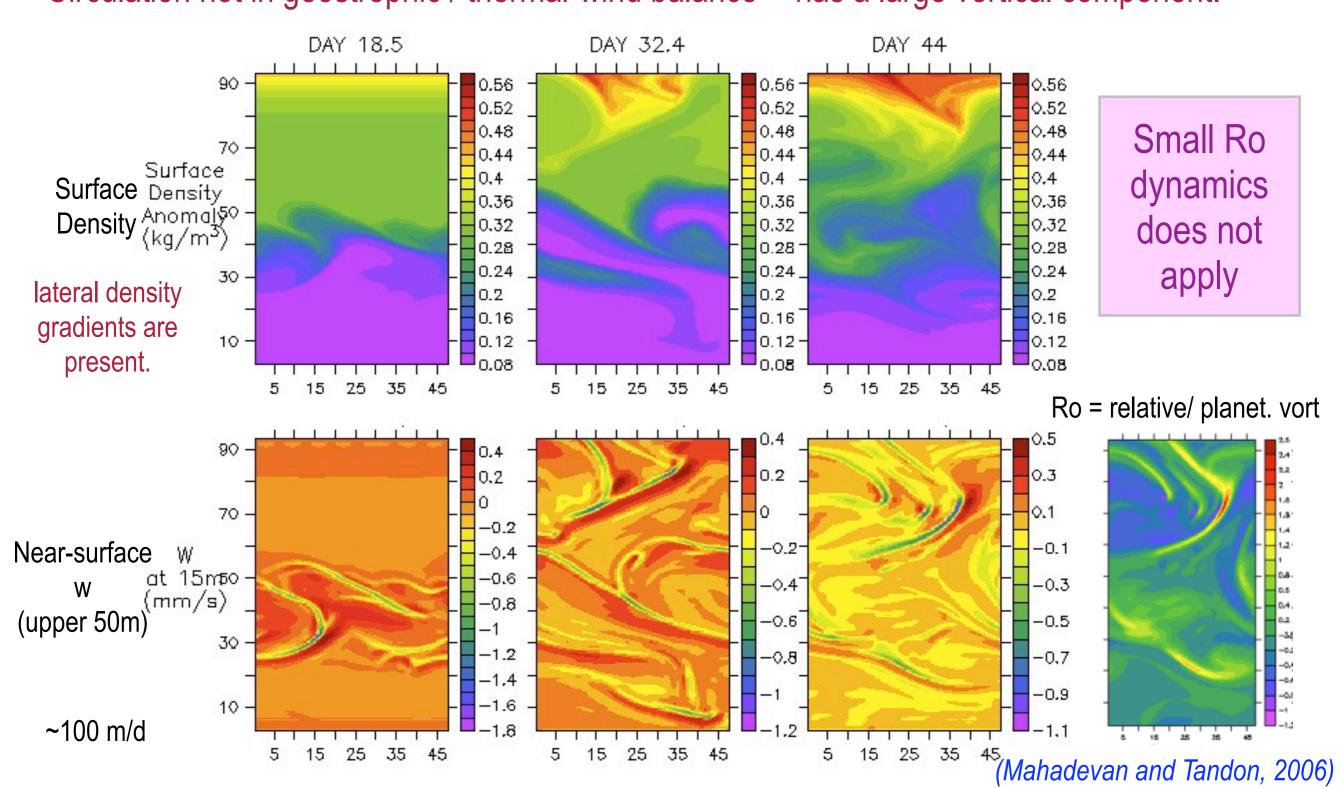
Solved efficiently using the multigrid method and line by line (block) relaxation.

Mahadevan et al., 1996a,b, Mahadevan & Archer 1998

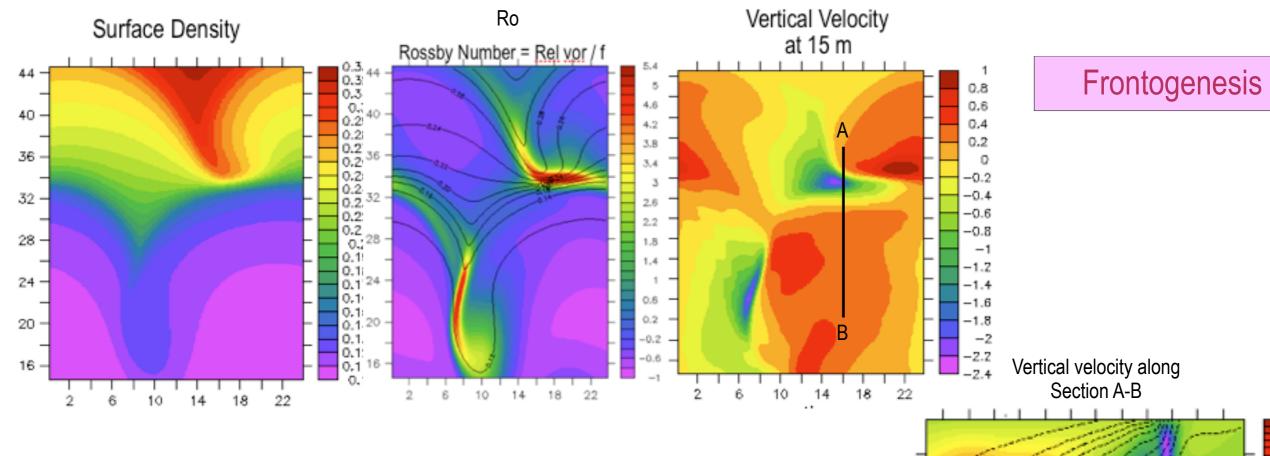
Strong vertical velocities

At higher (1 km) model resolution, we find that:

The largest vertical velocities O(100m/day) occur where the Rossby number becomes O(1). Circulation not in geostrophic / thermal-wind balance -- has a large vertical component.



A closer look at a single feature



A simpler model for circulation in the vertical plane

Semi-geostrophic: higher order in Ro

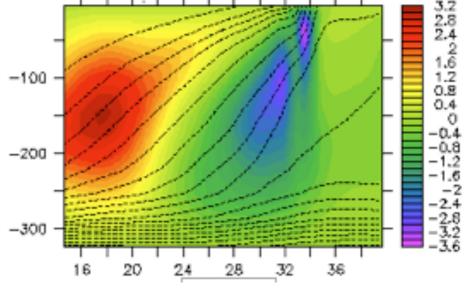
$$b = \frac{-g\rho;}{\rho_0}$$

$$F_2^2 \frac{\partial^2 \psi}{\partial z^2} + 2S_2^2 \frac{\partial^2 \psi}{\partial z \partial y} + N^2 \frac{\partial^2 \psi}{\partial y^2} = -2Q_2^g,$$

$$\text{where } N^2 = b_z, S_2^2 = -b_y = f u_{gz}, F_2^2 = f(f - u_{gy})$$

$$\text{Potential vorticity} = q_{2D} = \frac{1}{f} (F_2^2 N^2 - S_2^4)$$

$$\mathbf{Q}^g = (Q_1^g, Q_2^g) = \left(-\frac{\partial \mathbf{u}_g}{\partial x} \cdot \nabla b, -\frac{\partial \mathbf{u}_g}{\partial y} \cdot \nabla b \right)$$



< generally positive, but when it changes sign, this is not solvable

Loss of balance -- leads to vertical motion and mixing.

Summary

On large scales, the wind drives the ocean

Stratification and rotation inhibit vertical motion

Lateral dispersion is highly non-uniform

Fronts - generate large vertical velocities